

Astroparticle Physics & Cosmology

Summer term 2014

① Astroparticle Physics (G. Sigl)
→ Apr. 2 - May 21

② Cosmology (A. Westphal)
→ May 23 - July 11

Wednesday 10.30 am - 12 pm, Hörs. III

Friday 10.15 - 11.45 am Hörs. III

problem sessions: Friday 12.15 - 1.45 pm
Sem R. 1

course homepage, part A:

www.desy.de/~westphal

[/cosmology_2014/cosmology.html](http://www.desy.de/~westphal/cosmology_2014/cosmology.html)

↳ problem sets & lecture notes
as pdf's

please note:

↳ advanced theoretical course on
gravitation & cosmology:
winter term 14/15

23.5. Cosmology

literature:

Kolb & Turner: *The Early Universe*
(1990)

Mukhanov: *Physical Foundations of
Cosmology* (2005)

Weinberg: *Gravitation & Cosmology*
(1972)
Cosmology (2008)

Ryden: *Introduction to Cosmology*
(2003)

Dodelson: *Modern Cosmology* (2003)

Goenner: Einführung in die
Kosmologie (1996)

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⋮
Google/ArXiv: TASI Lectures:
Introduction to Cosmology
(2004) Trodden, Carroll

⋮

subject of the course (part A)³

particle physics:

microscopic structure of matter
distance scale: $\sim 10^{-19}$ m at LHC
corresponds via uncertainty principle
to an energy/temperature:

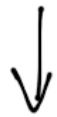
$$\lambda = \frac{\hbar}{p} \Rightarrow 1 \text{ fm} = 10^{-15} \text{ m} \approx \frac{\hbar c}{200 \text{ MeV}}$$

$$\sim 10^{-18} \dots 10^{-19} \text{ m} \hat{=}$$

$$E \sim 100 \dots 1000 \text{ GeV}$$

$$T \sim 10^{15} \dots 10^{16} \text{ K}$$

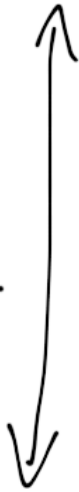
$$T = \frac{E}{k} \left(k = 8.5 \cdot 10^{-5} \frac{\text{eV}}{\text{K}} \right) \quad 4$$



results of particle physics (QFT) give description of hot plasma at temperature T

theoretical extrapolation up to $\sim 10^{16}$ GeV (neutrino physics, GUTs inflation)

LHC: Higgs mechanism (?), elementary scalar fields, QCD vacuum, cosmological phase transitions, supersymmetry? ...



... merges with the early universe: 5
 since E. Hubble's observations → the universe is expanding

$$v(r) = H_0 \cdot r$$

↑
 relative speed of recession of galaxies, supernovae, quasars ...

↑ distance of galaxies ...

2 plots

H_0 : present day Hubble parameter

$$H_0 = h_0 \cdot 100 \frac{\text{km}}{\text{s} \cdot \text{Mpc}}$$

$$h_0 = 0.71 \pm 0.06 \text{ HST Key Project}$$

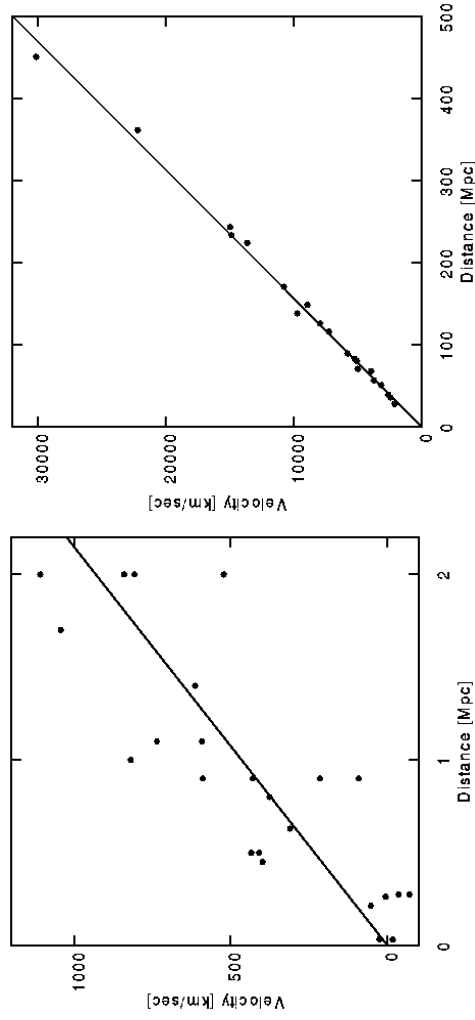


Figure 2.2: Hubble diagrams (as replotted in [17]) showing the relationship between recession velocities of distant galaxies and their distances. The left plot shows the original data of Hubble [18] (and a rather unconvincing straight-line fit through it). To reassure you, the right plot shows much more recent data [19], using significantly more distant galaxies (note difference in scale).

[17] Ned Wright's Cosmology Tutorial, <http://www.astro.ucla.edu/~wright/cosmo01.htm>

[18] E. Hubble, Proc. Natl. Acad. Sci., **15**, 168-173 (1929).

[19] A. G. Riess, W. H. Press and R. P. Kirshner, Astrophys. J. **473**, 88 (1996) [arXiv:astro-ph/9604143].

extrapolating back:

~ the earlier, the hotter...

We can list the early phases by their energy/temperature scale:

~ eV (atomic physics)

~ recombination of e^- and p into H \Rightarrow cosmic microwave background (CMB)

~ structure formation: stars, galaxies \leftrightarrow baryonic & dark matter

$\sim \text{MeV}$ (nuclear physics)

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\sim nucleosynthesis:

${}^1\text{H}, {}^4\text{He}, \text{D}, \text{T}, {}^3\text{He}$ from n, p, \dots

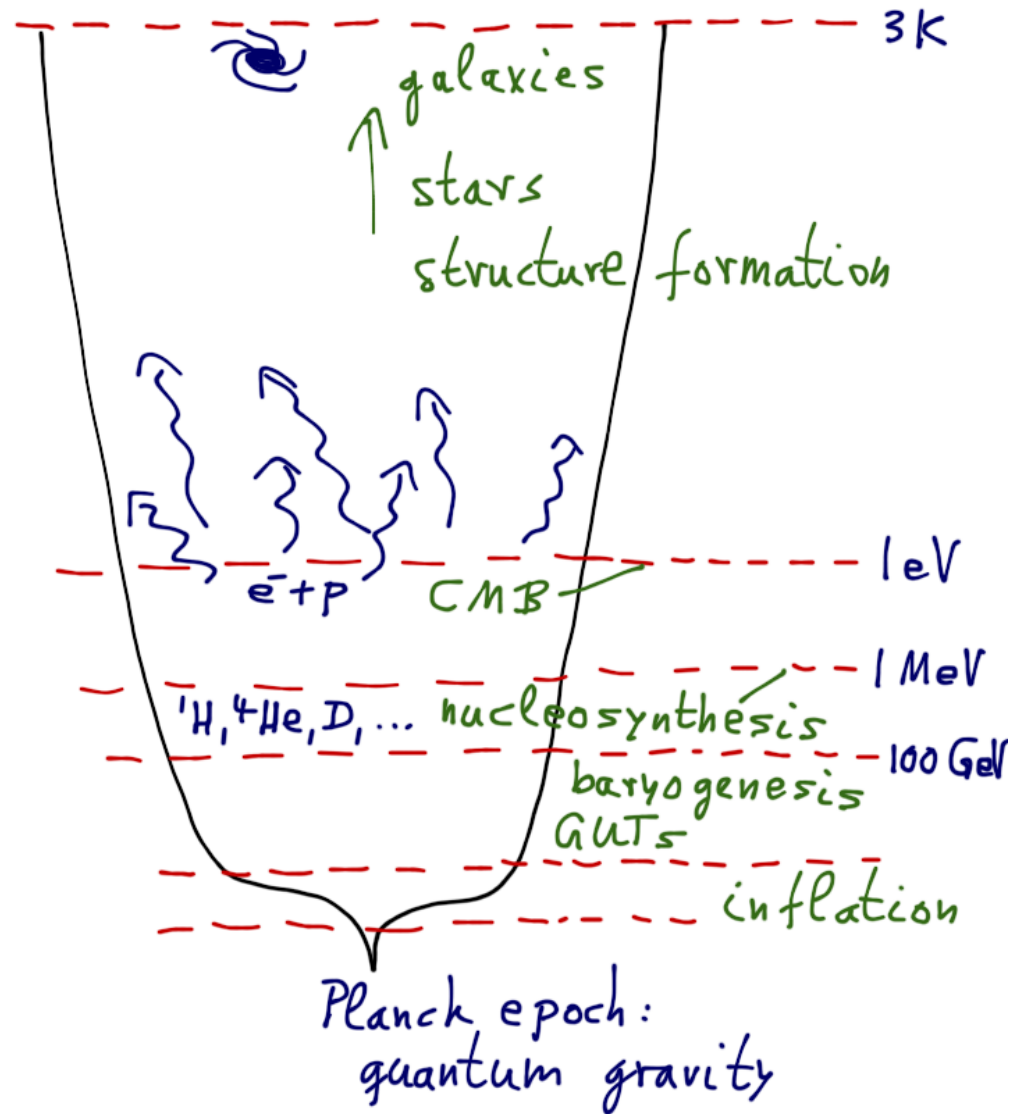
$\sim 100 \text{ GeV}$ (electroweak scale)

LHC: Higgs, Supersymmetry?

\rightarrow dark matter,
matter-antimatter/baryon asymmetry

:

? baryo-/leptogenesis,
inflation, ...



① The expanding universe ⁸

1. some cosmography:

units of distance:

$$\begin{aligned} \text{light year} = \text{ly} &= c \cdot \text{yr} \\ &= 3 \cdot 10^8 \frac{\text{m}}{\text{s}} \cdot \pi \cdot 10^7 \text{s} \\ &\simeq 10^{16} \text{ m} \end{aligned}$$

$$\text{parsec} = \text{pc} \simeq 3.26 \text{ ly}$$

distribution of (luminous & dark)
mass:

galaxies \rightarrow local groups \rightarrow
clusters/superclusters \rightarrow quasars

	$d [\text{ly}]$	⁹
center of the milky way	$2.8 \cdot 10^4$	
magellanic cloud	$2 \cdot 10^5$	
Andromeda galaxy (M31)	$2.1 \cdot 10^6$	
Virgo cluster	$6.6 \cdot 10^7$	
\vdots		
quasar	10^{10}	

(visible universe = CMB horizon
 \sim a few $\times 10^{10} \text{ ly} \simeq 10^{26} \text{ m}$)
(\rightarrow SDSS video)